





Overview of Synchrotron Radiation Research and the AOFSRR

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Asia/Oceania Forum for Synchrotron Radiation Research

Founded - 2006

AOFSRR Objective

The objective of the AOFSRR is to <u>encourage regional collaboration</u> in, and to <u>promote the advancement of, synchrotron radiation research</u> and related subjects in Asia and Oceania.

Specific Activities:

- (1) The annual workshop and Cheiron School, and organization of other scientific collaboration meetings;
- (2) Exchange of information of facilities and user groups;
- (3) Provision of a framework for cooperative activities;
- (4) Any activities that promote and expand the role of synchrotron light source facilities and synchrotron based research in the Asia Oceania region.



European Union



Large: ESRF (+PETRA)

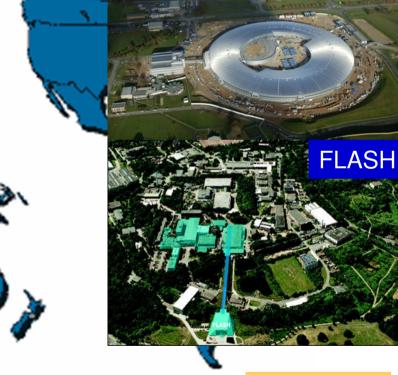
Medium: Diamond, Soleil, SLS

Soft X 3rg gen: Elettra, Max,

BESSY II ...

Next generation: FLASH,

European XFEL, FERMI, PSI..



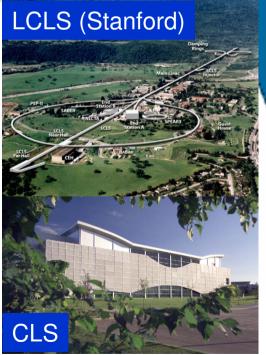


ESRF

Diamond

The Americas





Large: APS

Medium: CLS, NSLS II,

SSRL

Soft X 3rd Gen: ALS

2nd gen: NSLS, CHESS,

Brazil, Alladin...

Next generation: LCLS,

JLab, Cornell ERL(?)





Asia Oceania & the AOFSRR



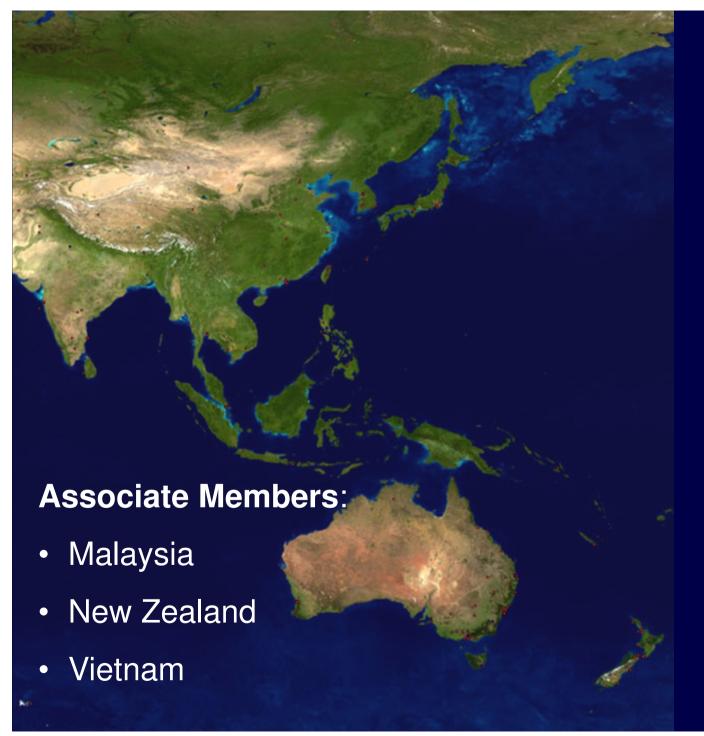
Large: SPring-8
Medium: AS, Indus II,
PLS II, Shanghai, TPS..
Soft X 3rd Gen: TLS,
UVSOR, Heifei..

2nd Gen: Beijing, PF.



- Facilities equal or better than Europe & USA
- Many bi-lateral agreements between facilities
- Few relationships between user communities
- No real regional organisation





Members:

- Australia
- China
- India
- Japan
- South Korea
- Singapore
- Taiwan
- Thailand





AOFSRR Activities





Annual Workshop

| Year | Host | | |
|---------|----------------------|--|--|
| 2006 | Tsukuba, Japan | | |
| 2007 | Hsinchu, Taiwan | | |
| 2008 | Melbourne, Australia | | |
| 2009 | Shanghai, China | | |
| 2010 | Pohang, South Korea | | |
| 2011/12 | Bangkok, Thailand | | |
| 2013 | Himeji, Japan | | |
| 2014 | Hsinchu, Taiwan | | |
| 2015 | Melbourne, Australia | | |

Cheiron School: Always SPring-8!!



User Community Networking

- The AOF annual workshop
- Cheiron School
- Open access to facilities
- Special Access for shutdowns/natural disasters
- Multi-nation scientific collaborations
- Regional accelerator school
- Other workshops



Promote Synchrotron Research

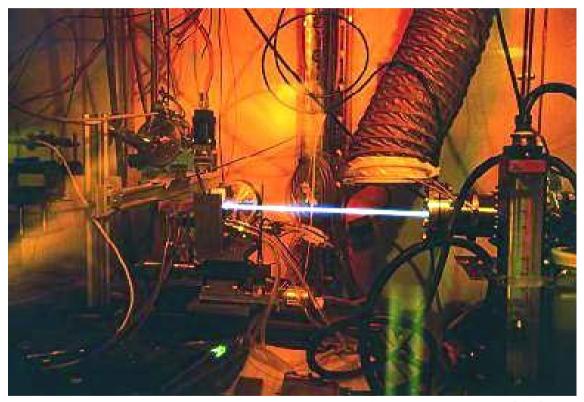
- Nations can build new communities at other facilities
 - Australian soft X-ray program at NSRRC and
 - NSRRC hard X-ray program at SPring-8
 - Indian beamline at the Photon Factory
- Promote SR research in non-member nations in the A-O region
- Assist SR science in developing nations
 - Cheiron School
 - Assistance to attend conferences
 - Work with other organisations (IUCr etc)



In the Future it is <u>your</u> AOF: How Should it Develop?



Synchrotron Radiation



X25 wiggler beam, NSLS

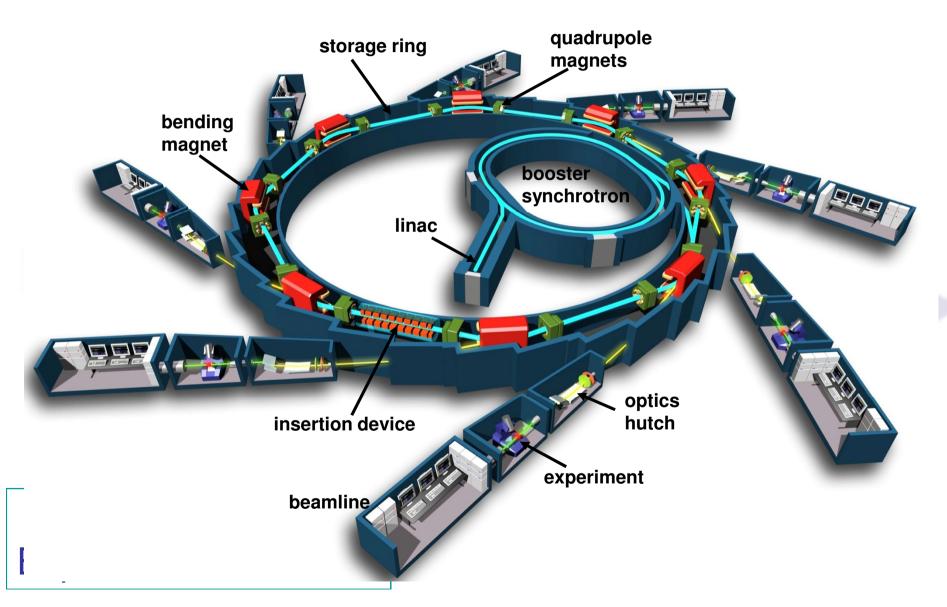


Outline

- What is a synchrotron?
- How is the light produced & what are its characteristics?
- Brief Basics of Synchrotron Beamlines
- Some Applications
- The Future (is here already): "Next Generation Sources"



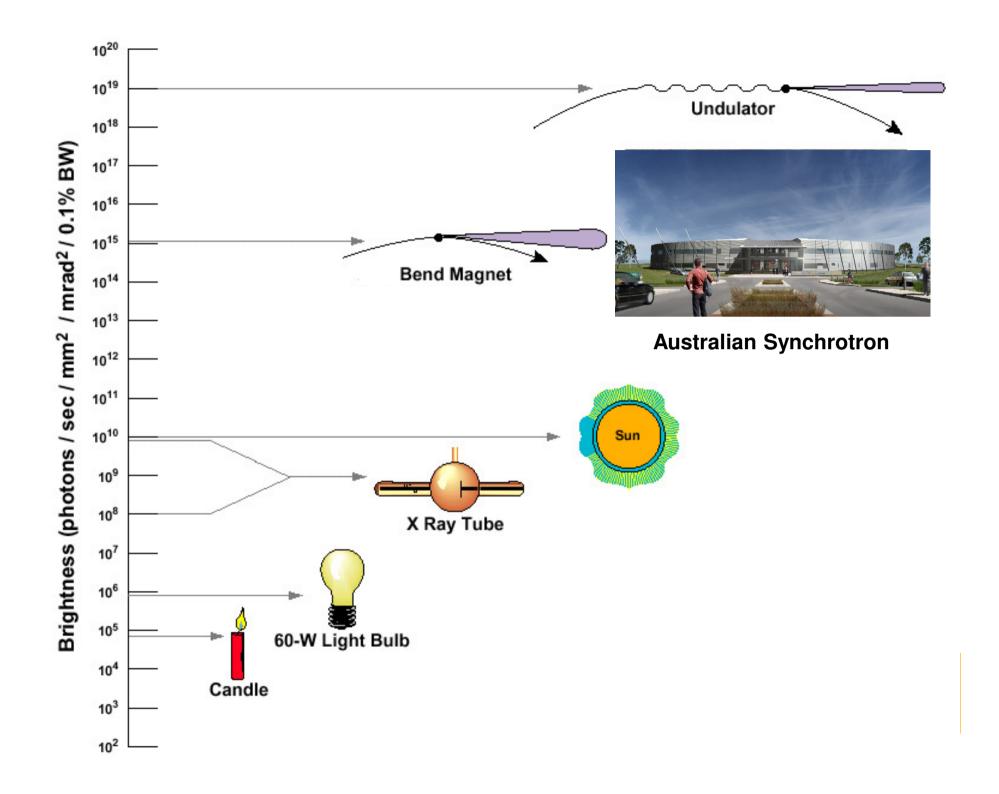
A Synchrotron Step by Step



Unique Characteristics of Synchrotron Radiation

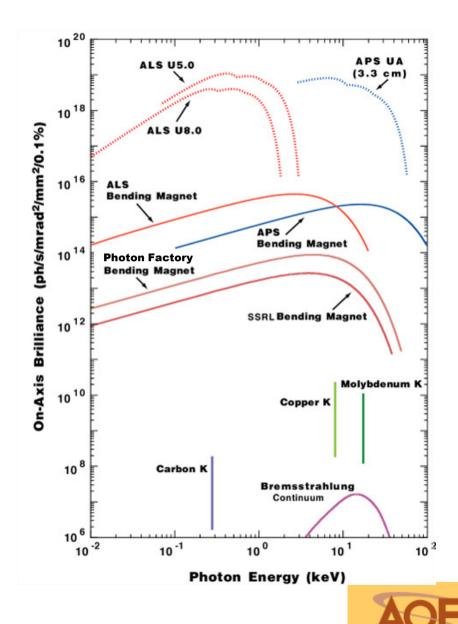
- Extremely high brightness. Modern synchrotron sources are about 10 billion times as intense as a laboratory X-ray generator: dilute samples; fast measurements; trace elements;
 - Low divergence: high intensity can be focussed onto tiny samples: Microscopies
- Wide X-ray energy spectrum:
 - the optimum X-ray energy to be chosen for each experiment;
 - X-ray spectroscopies are possible eg EXAFS
- Polarisation: various dichroisms; magnetic imaging; molecular orientation;
- Time structure: time of flight and very fast timing.



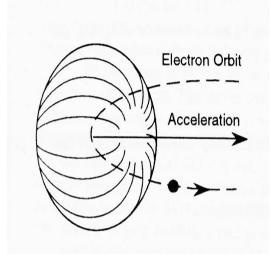


Characteristics of Synchrotron Radiation

- ✓ High brightness/flux
- ✓ Wide energy spectrum
- ✓ Plane polarised
- ✓ Pulsed



Generation of Synchrotron Radiation: Radiation from Accelerating Charge



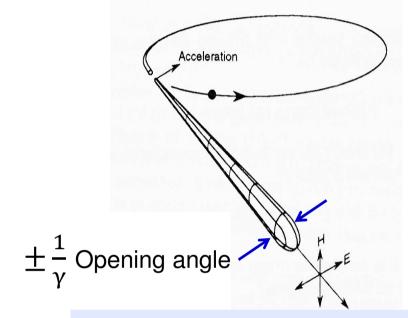
Low energy electrons OR electron frame:

Radiation in all directions

Example: Radio waves

from a transmitter.

$$rac{\mathbf{1}}{\gamma} = rac{\mathbf{m_0 c^2}}{\mathbf{E}} = \sqrt{\mathbf{1} - \left(rac{\mathbf{v}}{\mathbf{c}}
ight)^2}$$



High energy (relativistic) electrons – Laboratory frame:

Radiation pattern swept into a narrow cone in the forward direction = High brightness!

E = electron beam energy





Singapore Light Source

 $\gamma = 1400$

.7 mrad .04º

700 MeV



Australian Synchrotron

 $\gamma = 6000$

.2 mrad .01º

3 GeV



Spring-8

8 GeV

 $\gamma = 16000$

.06 mrad .004º

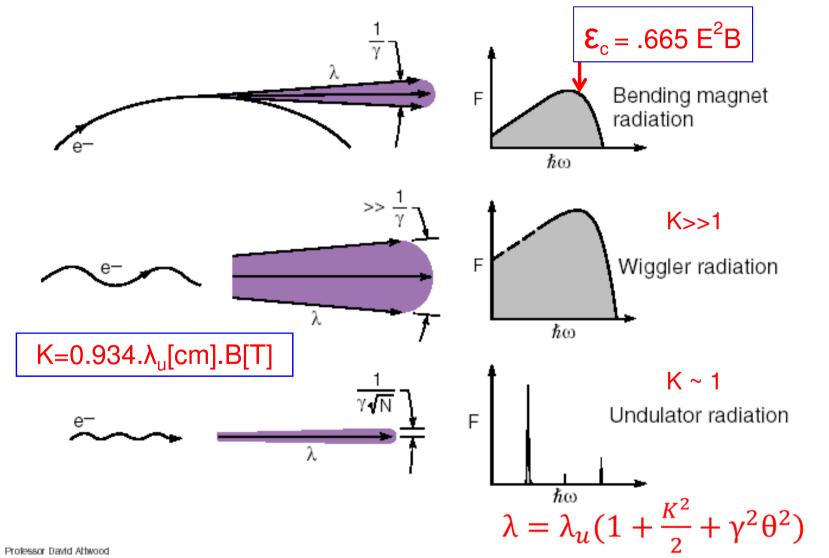


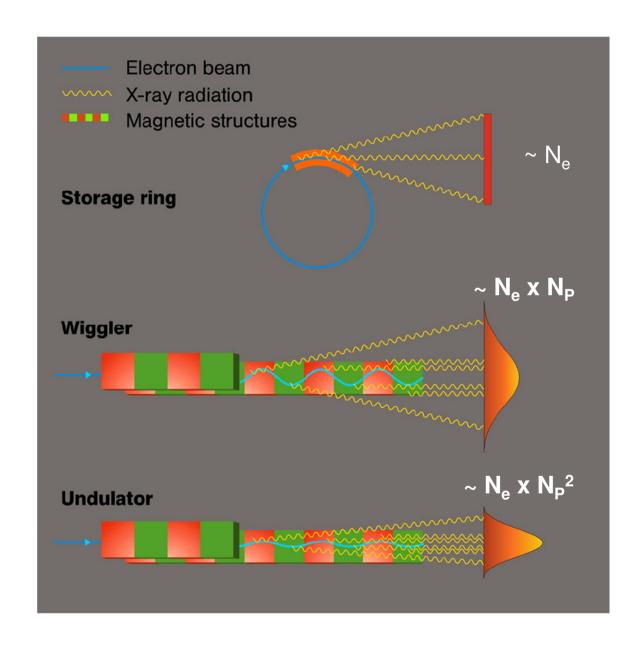
Third Generation Sources: Undulator Insertion Devices . Length@5 m -1st, 2nd Generation 3rd Generation: Insertion Devices Circular Photons electron motion Many straight sections containing Continuous periodic magnetic circular structures Photons trajectory · Tightly controlled electron beam Gap @ 1-3 cm **Electrons Bend Magnet** Undulator Radiation Radiation **Synchrotron** Radiation X-ray X-rays light bulb · Laser-like Photon flux Photons/s Tunable 600x10¹² 500 -500 eV Photon energy Photon energy Phot/s/0.1%bw 400 -300 -200 100 -2000 4000 6000 8000 10000eV Photon Energy



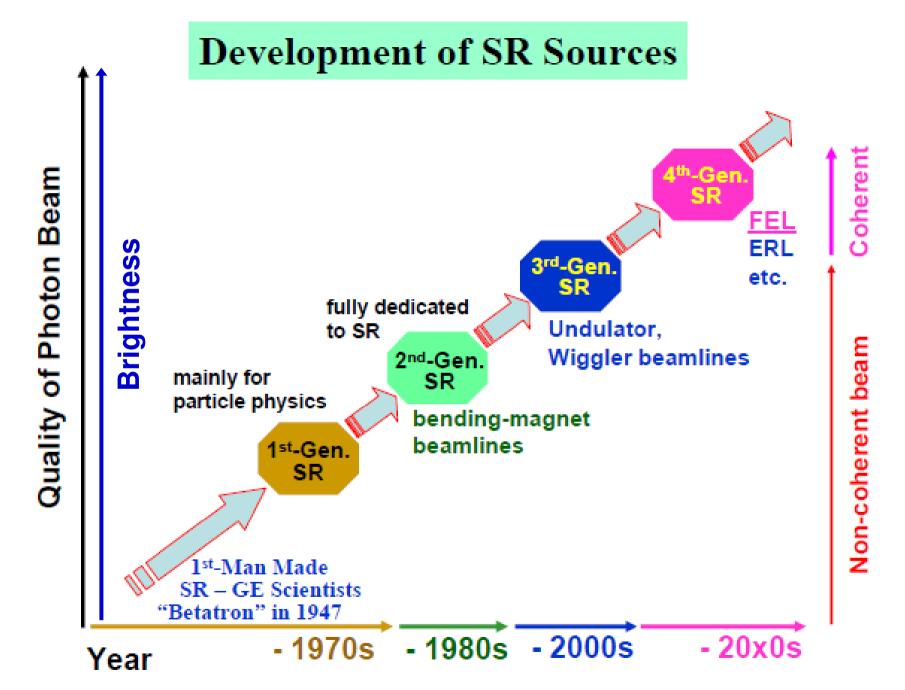
Three Forms of Synchrotron Radiation









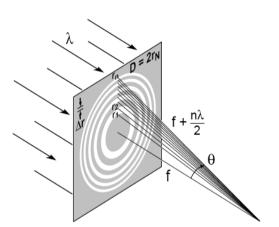


SR Development is towards higher & higher brightness – why?

- Need high spatial resolution:
 - Tiny samples (protein crystals...)
 - Nano-structure materials
 - Non-homogeneous samples...
- Focusing optics image the source
 - Smaller focus (smaller and smaller xtals) or
 - More working dist for same focus size
- New imaging techniques especially coherence based imaging
- etc

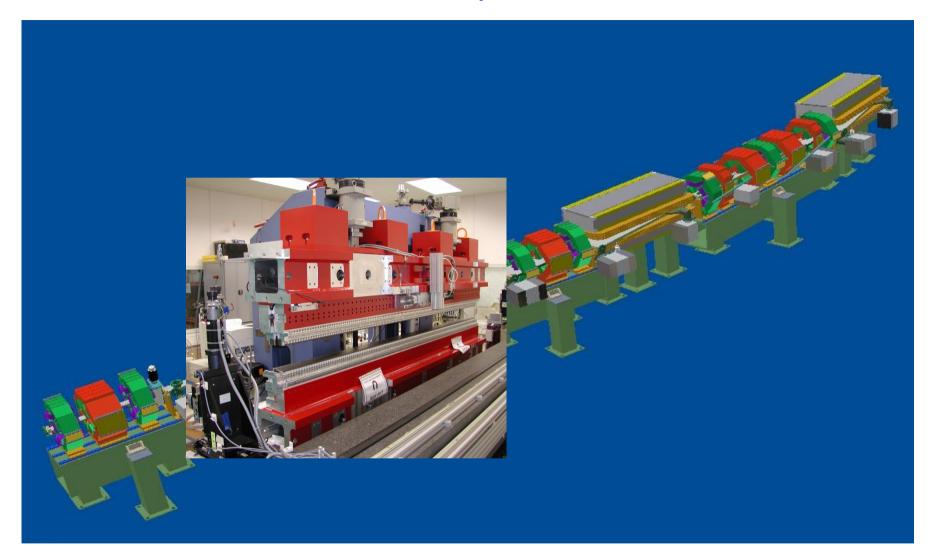








Section of the Australian Synchrotron





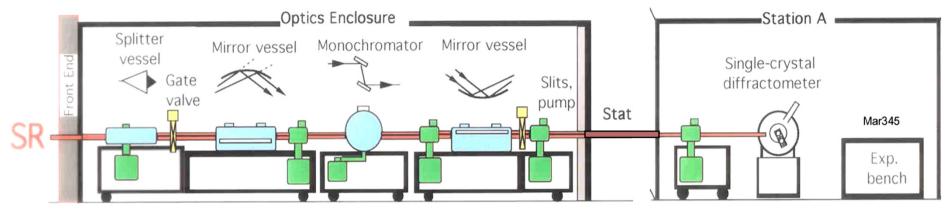
Beamline Design Goals

- Deliver the required X-ray beam to the experiment:
 - Energy and bandwidth
 - Spot size
 - Divergence/convergence
- Preserve source characteristics eg intensity, brightness, coherence
- Optimise signal / background
- Be very stable and reproducible, in position, intensity and energy
- Be safe to operate
- Be user friendly to operate
- Achieve all the above within a reasonable budget!

(Good Luck!)

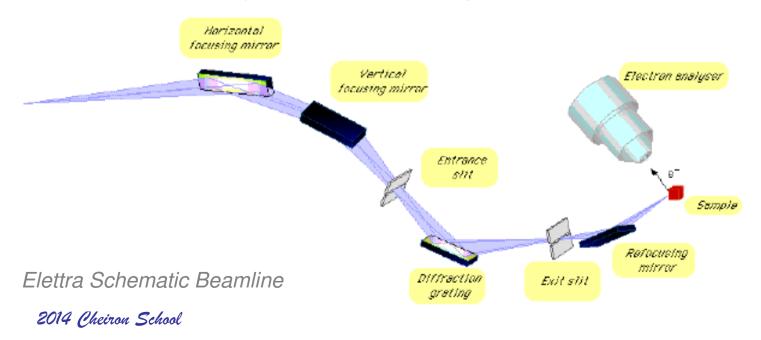


Hard X-ray Beamline: Si crystal monochromator E > 4 keV



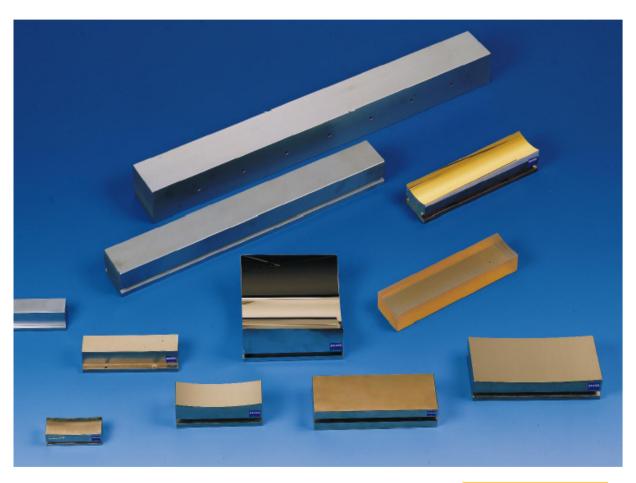
(Swiss-Norwegian beamline, ESRF)

Soft X-ray Beamline: Grating monochromator E < 2 keV



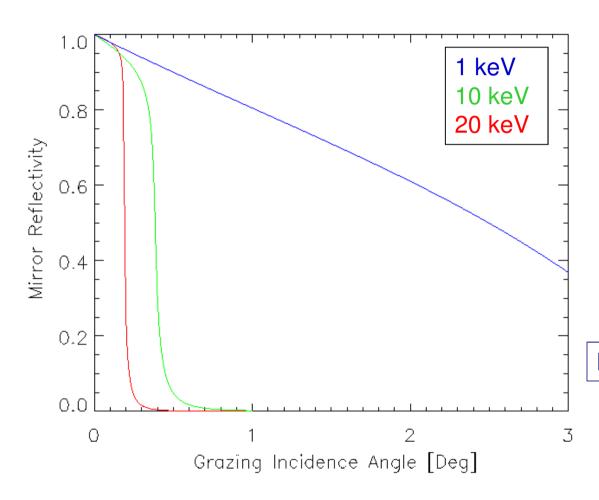
Mirrors for Synchrotron Beamlines

- Deflection
- Focusing
- Harmonic Rejection
- Power Reduction





Critical Angle/Reflectivity with Energy: Rhodium Coated Mirror Example



Harder X-rays need more grazing angles and longer mirrors:

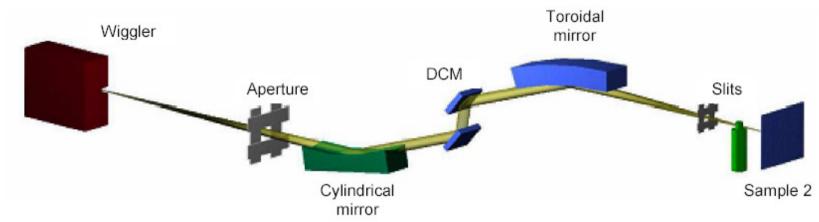
2 mm high beam needs:

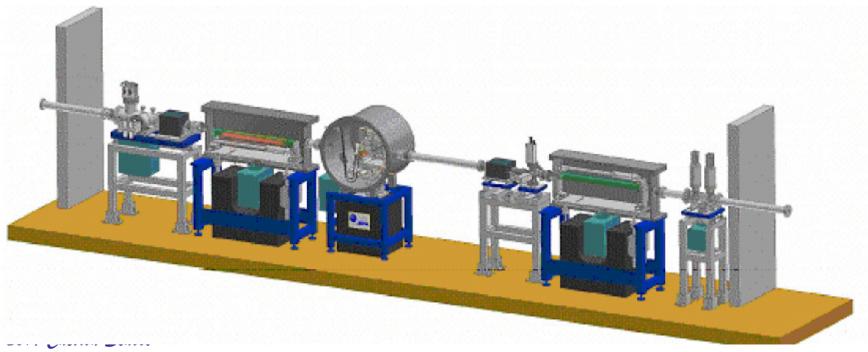
- ≤ 10 cm mirror at 1 keV
- ≥ 80 cm mirror at 20 keV

Mirrors are a low pass filter

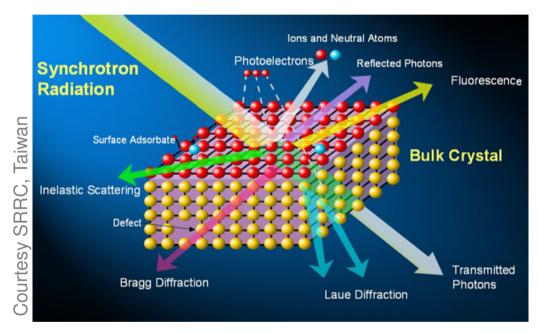


An example beamline: the AS Xray Absorption Spectroscopy Beamline





X-rays and their Interaction with Matter



X-ray Diffraction → Structure

X-ray Fluorescence

→ trace element analysis

Transmitted Photons:

→ Imaging Absorption Spectroscopy

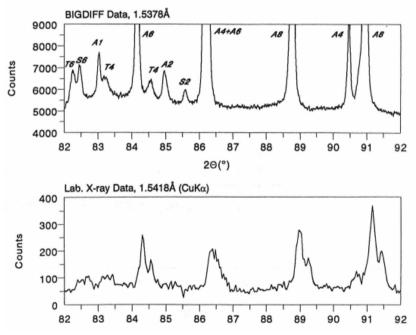
→ Chemical information



| | Synchrotron | Proton | Electron Microscope | SIMS | Neutron |
|-------------------------|-------------|--------|------------------------|----------|----------|
| Sensitivity | ✓ | • | * | ✓ | × |
| Sub micron | √ | • | √ | * | × |
| Chemical Information | √ | * | | | × |
| In-situ | √ | • | * | * | ✓ |
| Atomic Structure | ✓ | × | √ | * | ✓ |



Sometimes High Intensity = Better Data Synchrotron Powder XRD



Multi-phase ceramic: α Al₂O₃, ZrO₂, MgO-Al₂O₃ (spinel). Top synchrotron data; Bottom: lab data.

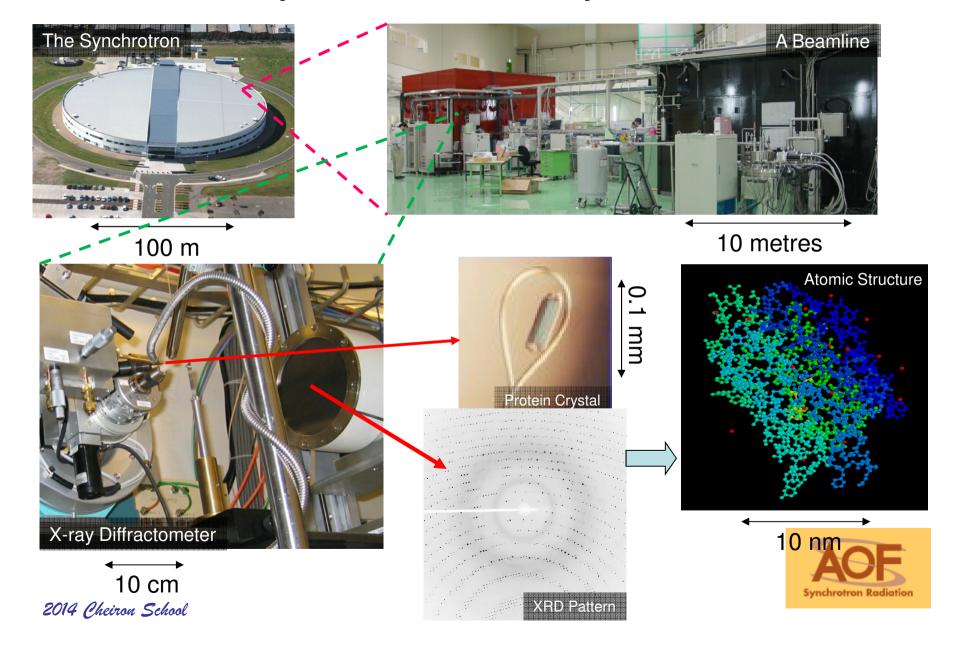
B. H. Oconnor, A. van Riessen, J. Carter, G. Burton,

R. F. Garrett and D. J. Cookson,

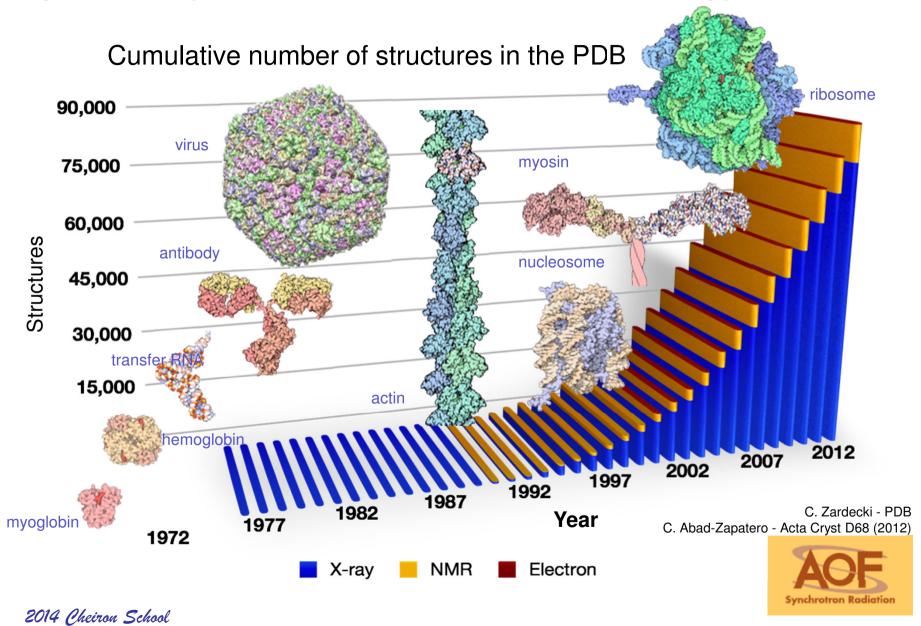
J. American Chemical Soc. 80 (1997) 1373



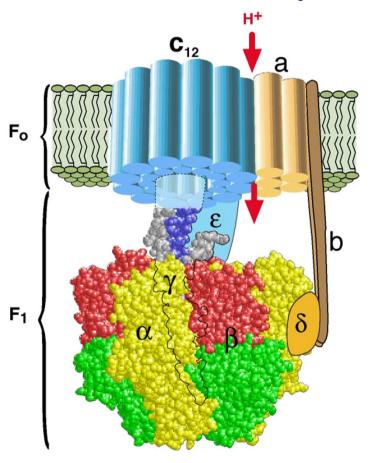
X-ray Diffraction at a Synchrotron



Impact of Synchrotrons on Structural Biology

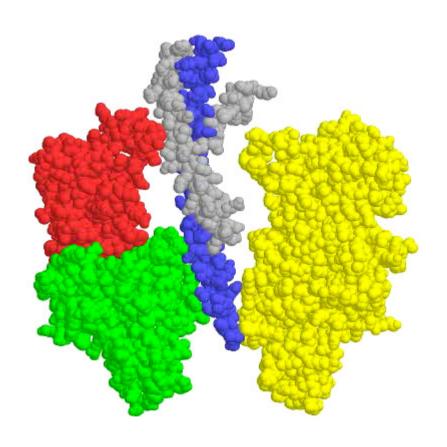


ATP Synthase: a Molecular Motor



H. Wang and G. Oster (1998). Nature 396:279-282.

John Walker won the 1997 Nobel Chemistry prize for solving the F1 catalytic domain using synchrotron radiation at Daresbury, UK.



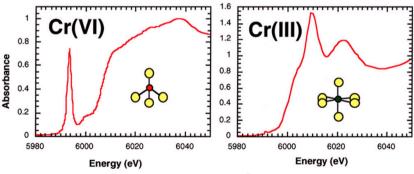
Atomic structure informs biological function

Broad Energy Spectrum: SR Only Spectroscopies eg Xray Absorption Spectroscopy

XANES: near edge structure

Sensitive to chemical environment of absorbing element.

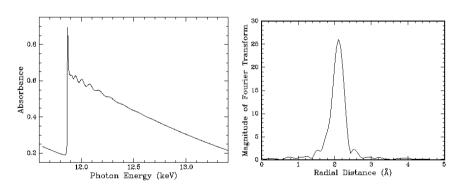
Often different valence states have markedly different XANES spectra.



XANES spectra of Cr III (relatively benign) and Cr VI, a known carcinogen.

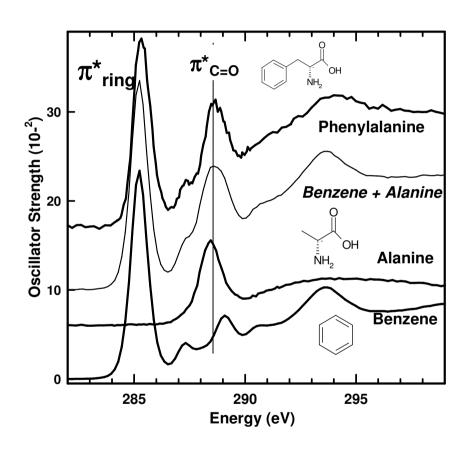
EXAFS: extended structure to ~1 keV above an absorption edge

Nearest neighbour atomic distances, coordination etc. Crystals not required: disordered systems like solution species can be measured.

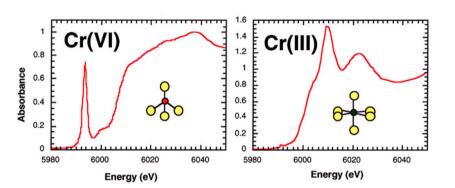


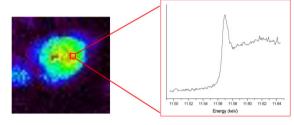
Amorphous GaAs EXAFS and Fourier transform.

Near Edge Spectroscopy: Chemical Sensitivity



Carbon K-edge Spectra

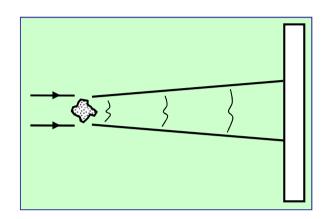




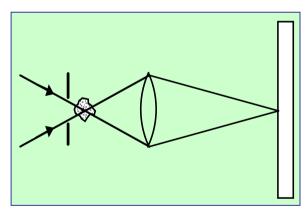
XANES of Pt located in a tumour cell Hambley, U Syd



Imaging

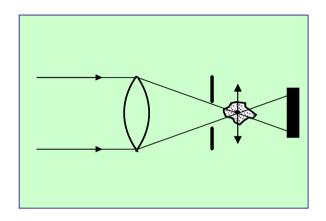


Radiography



Imaging

Mapping



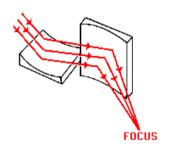
Scanning

Various Imaging Techniques have become more and more important in synchrotron research over the last 10 years



Some Imaging Needs Focusing optics

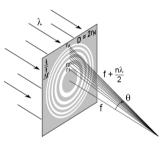
Reflective (Kirkpatrick-Baez mirrors) typical \sim 1 μ m High efficiency, achromatic, limited to \sim 10 nm



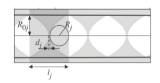
Diffractive (Fresnel zone plates)

Moderate efficiency, limited to ~10 nm

typical ~ 100 nm



Refractive (compound refractive lenses) 10s μm - \sim 50 nm Low efficiency, highly chromatic, aberrations Works well with high energy X-rays





Contrast mechanisms in x-ray imaging

> **Absorption** measure electron density; can be element specific

Fluorescence measure elemental distribution

Spectroscopy extract chemical state, spin state

Diffraction reveal structure, strain, magnetism, charge...

Phase measure real part of refractive index

In general with X-rays:

- Natural sample contrast is often possible; staining not required
- Image structure of thick samples, sectioning not required
- More penetrating, less damage, less charging than with electrons
- In situ imaging image samples in natural environment.



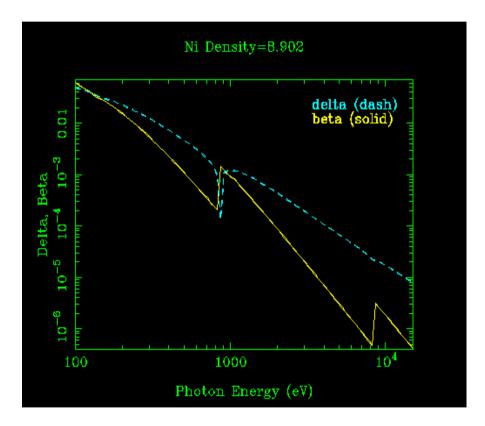
Phase Contrast

Refractive index:

for X-rays it is less than 1 by about 1 part in a million

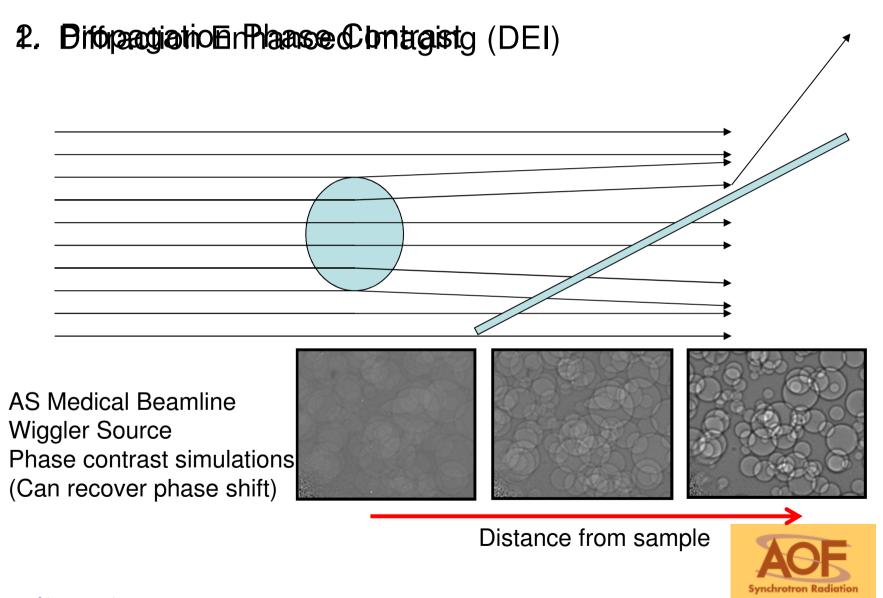
$$n = 1 - \delta - i\beta = 1 - \frac{r_e}{2\pi} \lambda^2 \sum_{i} n_i f_i(0)$$

- Absorption contrast: sensitive to *Im(n)*
- Phase contrast: sensitive to Re(n)
- At high X-ray energies, phase contrast wins



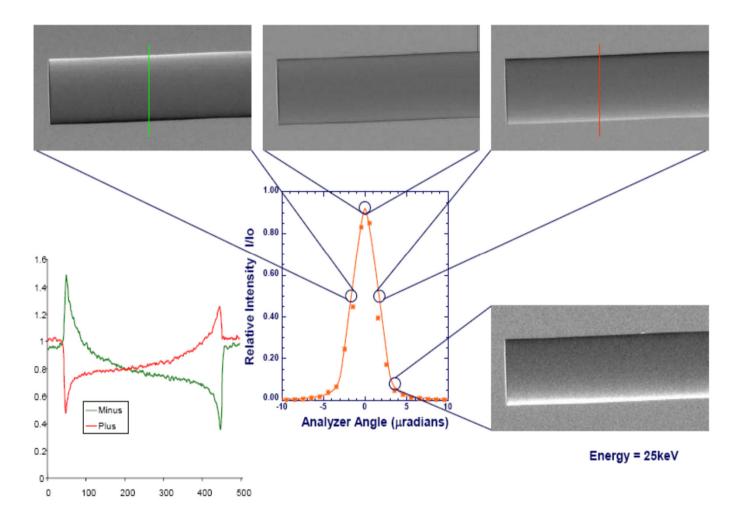


Phase Contrast



Diffraction Enhanced Imaging

- Edge/density gradient sensitive
- Move on rocking curve to change contrast





X-ray phase imaging: Biology and Materials



Conventional X-ray

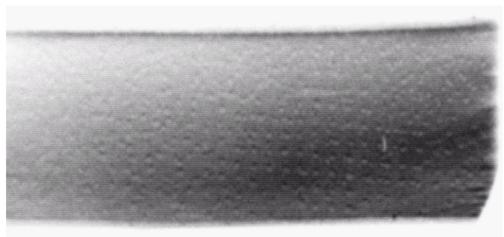


Synchrotron



Rob Lewis, Monash University/ Energy = 20keV

Synchrotron: propagation phase contrast

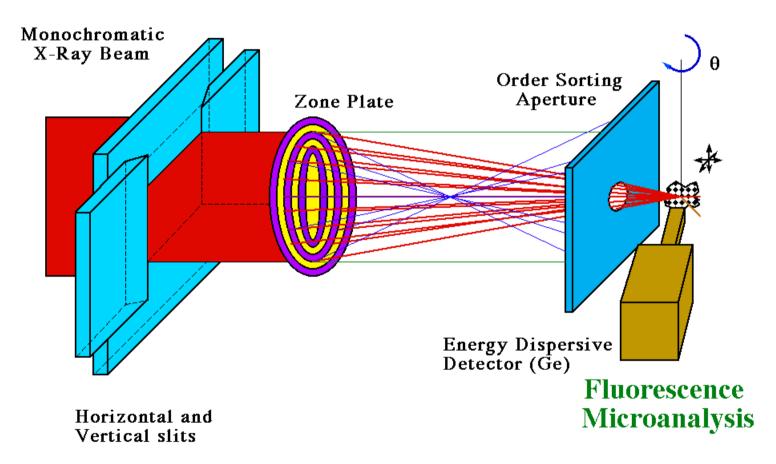


DEI "Dark field" phase image of bonded aluminium sheets @ 33 keV
Dots are bubbles in the epoxy bond.

Stevenson, Garrett, Hyodo et.al.



X-ray Fluorescence Microscopy

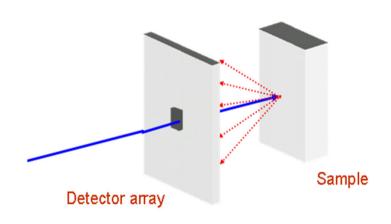


Conventional Fluorescence Microscope: APS 2ID-D

Advanced fluorescence detector at the AS

Annular geometry

- Maximises solid angle, sample @ 90°
- 384 Si pixel detector array (BNL, Siddons et al)
- No constraint on lateral sample size and scan range



solid-angle ~1.5-2 steradians

+ Parallel data processing

- CSIRO: HYMOD2 pipelined, parallel processor (Ryan etal)
- Whole XRF spectrum acquired and analysed in real time

+ Fast Scanning Stage

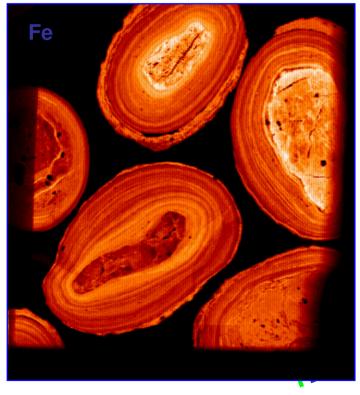
- Data acquired "on the fly"
- milli-second dwell times
 cf 1 second or greater normally

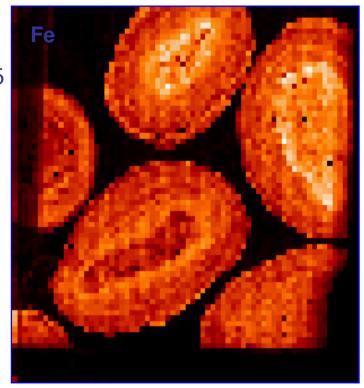
= New micro-XRF capability at the AS X-ray Fluorescence Microscope beamline

XFM image definition (number of pixels) limited by dwell time

Long dwell → **Low Image Definition**

- ~1 s / pixel (for readout of 1-16 detector spectra)
- 1.3 hours \rightarrow 67 x 67 pixels





Synchrotron Radiation

Short dwell → **Good image definition**

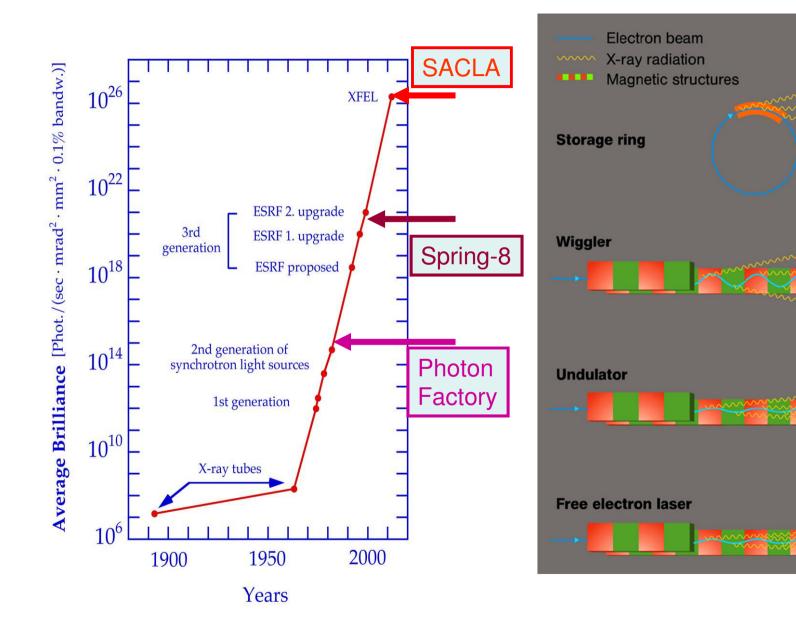
- 32 ms / pixel
- 1.3 hours → 375 x 375 pixels (30 fold increase)
- New Maia-32 prototype detector, NSLS X27A

New Sources:

XFELs



Next Step - X-ray Lasers? Yes → FELs



 $\sim N_e \times N_P$

 $\sim N_e \times N_P^2$

 $X N_P^2$

Light Source Performance XFEL 100 Degree of Coherence (%) Pulse Width (5ec)
10-14
Pulse Width (5ec)
10-12 ×₃10 10 10-15 0.1 10-11 SPring-8 10-10 0.0 × 10-**Photon Factory** 10¹⁰ 10. (Photons/sec/mm²/mrad²/0.1%b.w.) × 10 1030

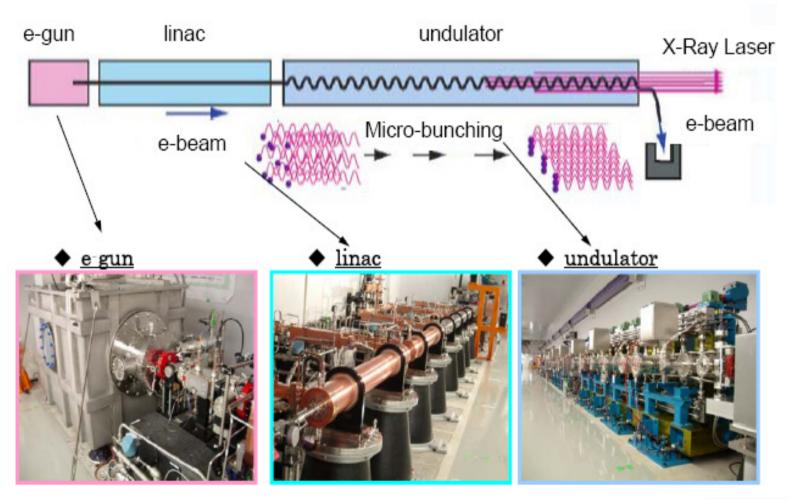
Remarkable Features of XFEL producing λ<0.1 nm X-Rays

- Migh Peak Brilliance
- Narrow Pulse Width
- High Degree of Coherence

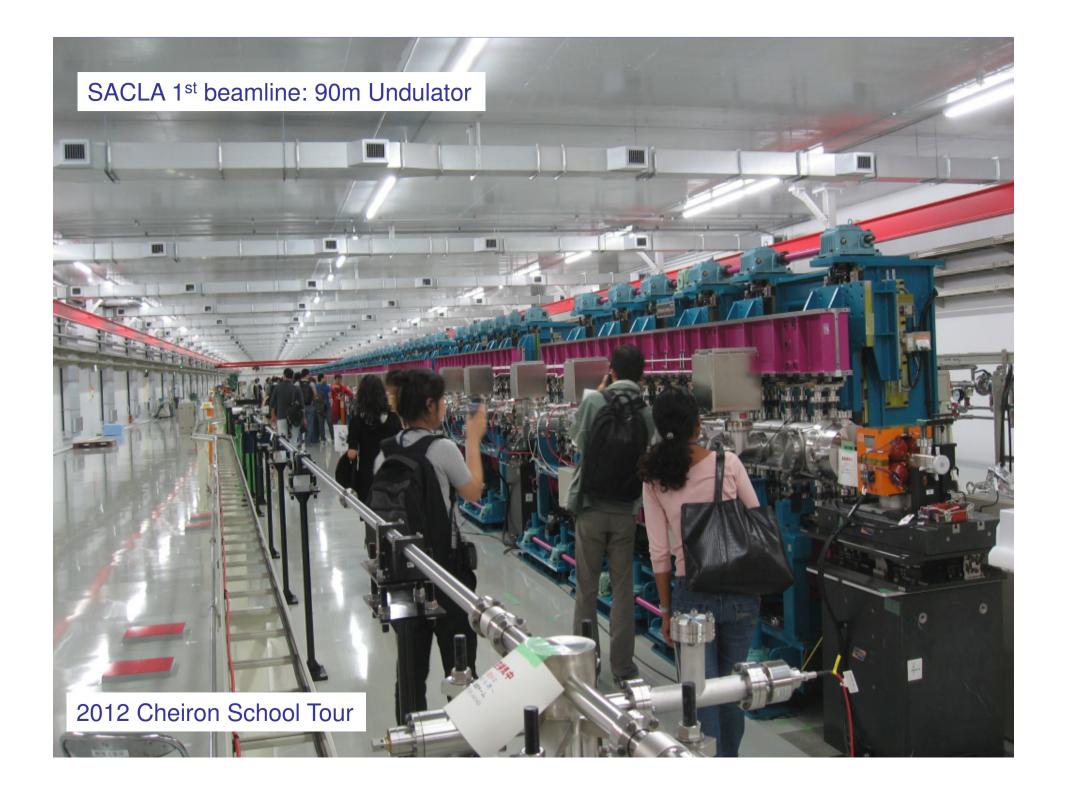
 4 Cheiron School



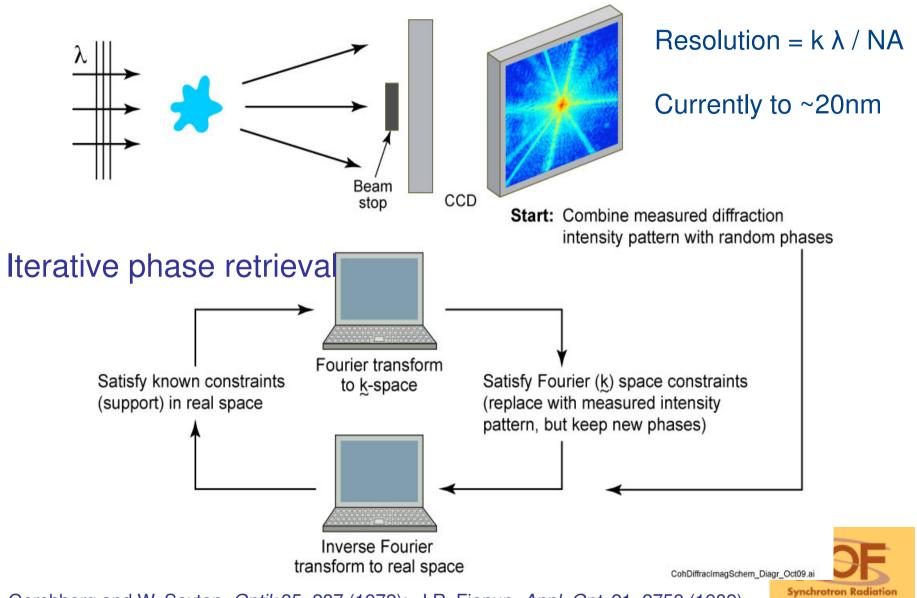
Linac-Based Free Electron Laser Self-Amplified Spontaneous Emission (SASE)



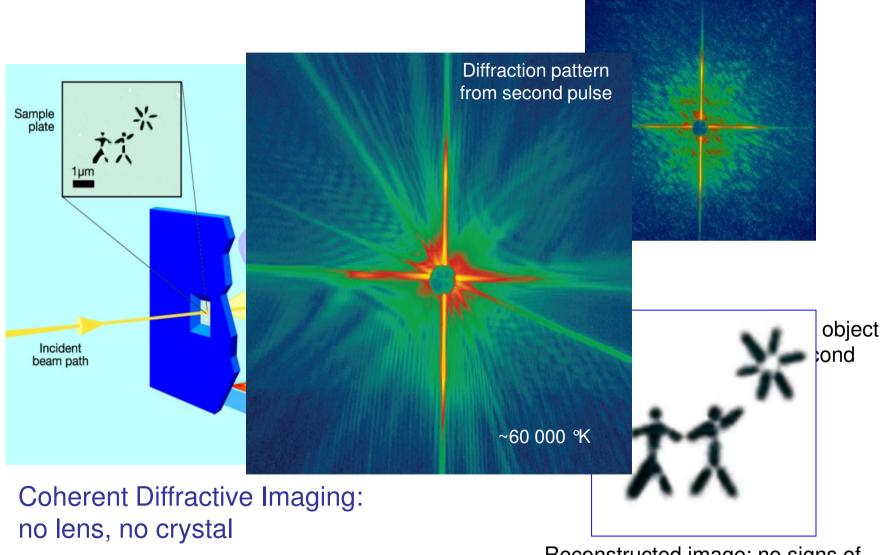




Coherent Diffractive Imaging: no lens, no crystal



"Diffract and Destroy": Single Shot Imaging at the FLASH Soft X-ray FEL



H. Chapman et al., Nature Physics 2, 839 (2006)

Reconstructed image: no signs of damage caused by the pulse.

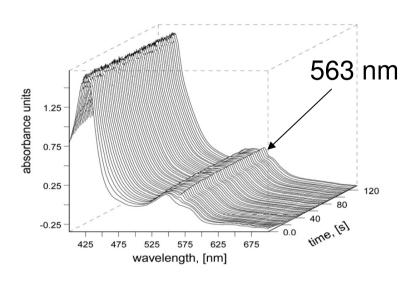
First serial femtosecond crystallography experiments at LCLS/AMO/CAMP - 7 A resolution

Chapman et al Gas focussed liquid jet: Nature 470: 73 (2011) 4μm diam., flow rate 10- 14 μl min, 10 ms/s De Ponte et al., J. Phys. D 41, 195505 (2008) Liquid jet λ =6.2-6.9 Å, 1.8-2.0 keV Rear pnCCD 10-300 fs pulse duration LCLS X-ray pulses (z = 564 mm)30 (60) Hz 10¹² photons/pulse Interaction Front pnCCD $(z = 68 \text{ mm})_{1800 (3600) \text{ patterns/min}}$ 900 J/cm² point Dose/pulse: 300-700 MGy ~ 7 µm diam.

~ 4 µm diam.

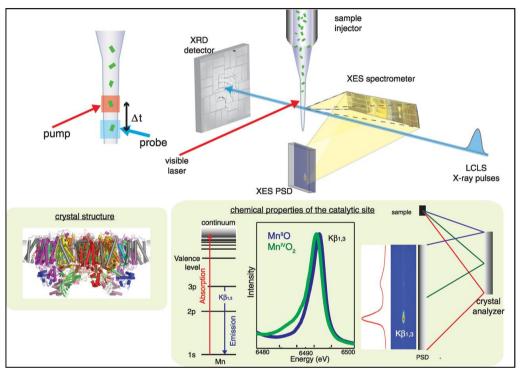
5 TB in one night

Radiation damage-free data collection



Beitlich et al., JSR (2007)

→ on-line spectroscopy
shows radiation damage in
chloroperoxidase crystal at a
synchrotron in seconds at a
low dose rate (28 kGy/s)



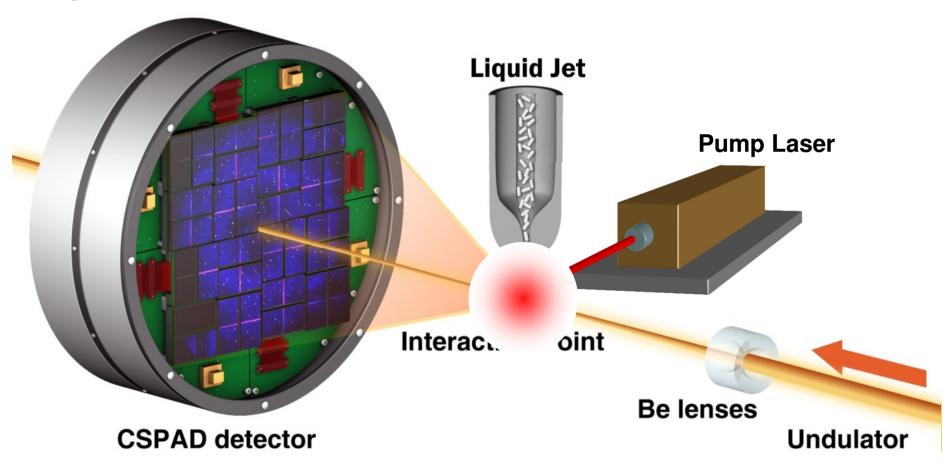
Kern et al., Science (2013)

→No radiation damage to metal cluster in Photosystem II during diffraction at LCLS

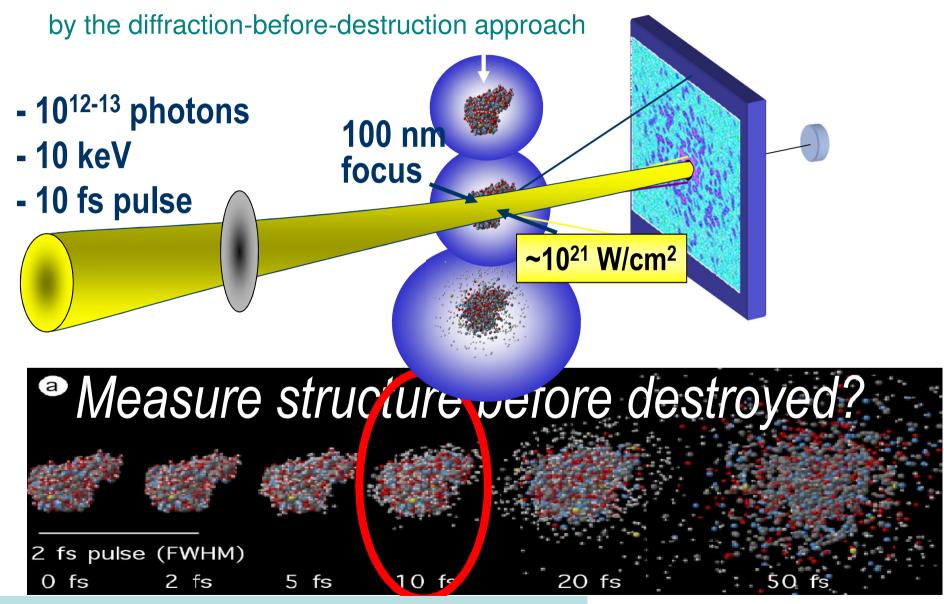
Synchrotron Radiation

Time-resolved pump-probe experiments on photosensitive proteins

Best time-resolved MX (Laue, synchrotron) done by F. Schotte/Ph. Anfinrud has 100 ps time resolution. With FELs this can be reduced to ps/hundreds of fs.



Coherent diffractive imaging of single particles



Calculations. in vacuum Neutze et al., Nature 2000

FELs: Many research areas

Femtosecond experiments

- pump-probe experiments on atoms and molecules (femto-chemistry)
- sum-frequency generation
- serial crystallography

Interaction of ultra-intense XUV pulses with matter

- multiphoton excitation of atoms, molecules, clusters...
- creation and characterizaton of dense plasmas
- imaging of nano-objects and biological samples

Investigation of extremely dilute samples

- photodissociation of molecular ions
- highly charged ions
- mass selected clusters

Investigation of surfaces and solids

- XUV laser desorption
- surface dynamics
- femto-magnetism
- study of highly correlated materials
- luminescence under FEL radiation
- meV-resolution photon and photoelectron spectroscopy of surfaces and solids with nm resolution

FELs are Great, BUT...

- Only a few of them and only a few (at the moment one) experiments at a time
 - Capacity very limited; difficult to access
 - Each experiment VERY expensive
- Technical challenges at every step
 - Source, optics, samples detectors, data etc
- SASE Process every pulse different
- Can we make a brighter synchrotron light source to "fill the gap"?

Brightness Limit is a Diffraction Limited Source

Emittance characterises the brightness of a source; it is the product of source size and divergence in a particular dimension.

For SR from a single electron, the emittance/brightness is diffraction limited:

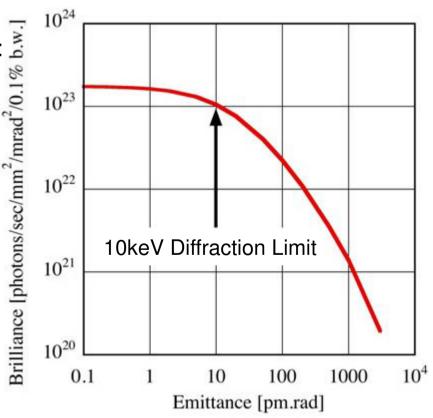
$$\epsilon_r = \sigma_r \sigma_{r'} = \lambda/4\pi$$

So wavelength corresponding to the diffraction limit for a particular emittance:

$$\lambda_{\mathbf{DL}(\mathbf{x},\mathbf{y})} = 4\pi\epsilon_{\mathbf{x},\mathbf{y}}$$

Real emittance is combination of radiation emittance and real source size/divergence:

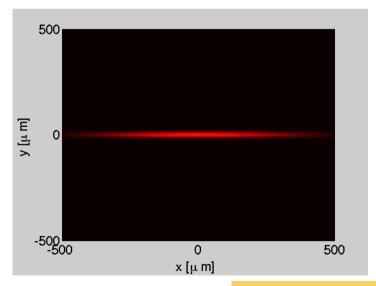
$$egin{aligned} oldsymbol{\Sigma_{x,y}} &= \sqrt{\sigma_{x,y}^2 + \sigma_{r}^2} \ oldsymbol{\Sigma_{x',y'}} &= \sqrt{\sigma_{x',y'}^2 + \sigma_{r'}^2} \end{aligned}$$



Brightest 3rd Generation Storage Ring Sources:

| Source | energy GeV | ϵ_x nmrad | max. B_n 10^{20} | circ. m | $egin{array}{c} \lambda_{Dlx} \ 	ext{Å} \end{array}$ | lattice |
|----------|---------------|--------------------|----------------------|------------|--|---------|
| ESRF | 6 | 4 | 3.0 | 844 | 500 | DBA |
| APS | 7 | 3.1 | 1.4 | 1104 | 390 | DBA |
| SPring-8 | 8 | 2.7 | 2.5 | 1436 | 340 | DBA |
| DIAMOND | 3 | 2.7 | 1.5 | 561 | 340 | DBA |

- Horizontal emittance 3-20 nm.rad
- Vertical emittance ~100 times less
- Very asymmetric beam in straight sections:
 - ➤ Horizontal ~ 1mm FWHM
 - ➤ Vertical ~ 10 µm
- Far from diffraction limit in Horizontal
 @ 10keV
- Close to diffraction limit in vertical





How to Design a Brighter Ring?

The horizontal emittance is a characteristic equilibrium quantity for each storage ring

$$\epsilon_{\mathbf{x}} \propto \mathbf{E^2} \cdot \mathbf{\theta^3} \cdot \mathbf{\Gamma}$$

angle per bending magnet: heta lattice dependent quantity: $oldsymbol{\Gamma}$ electron energy: $oldsymbol{\mathsf{E}}$

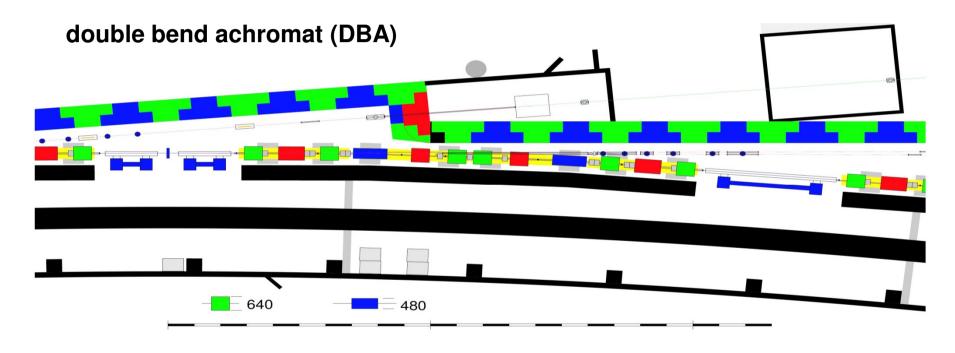
Emission of SR also reduces emittance – "natural emittance" of a storage ring is equilibrium between SR and energy spread.

Reduce energy (not too much or no X-rays)
Reduce angle of bending magnets
More synchrotron radiation



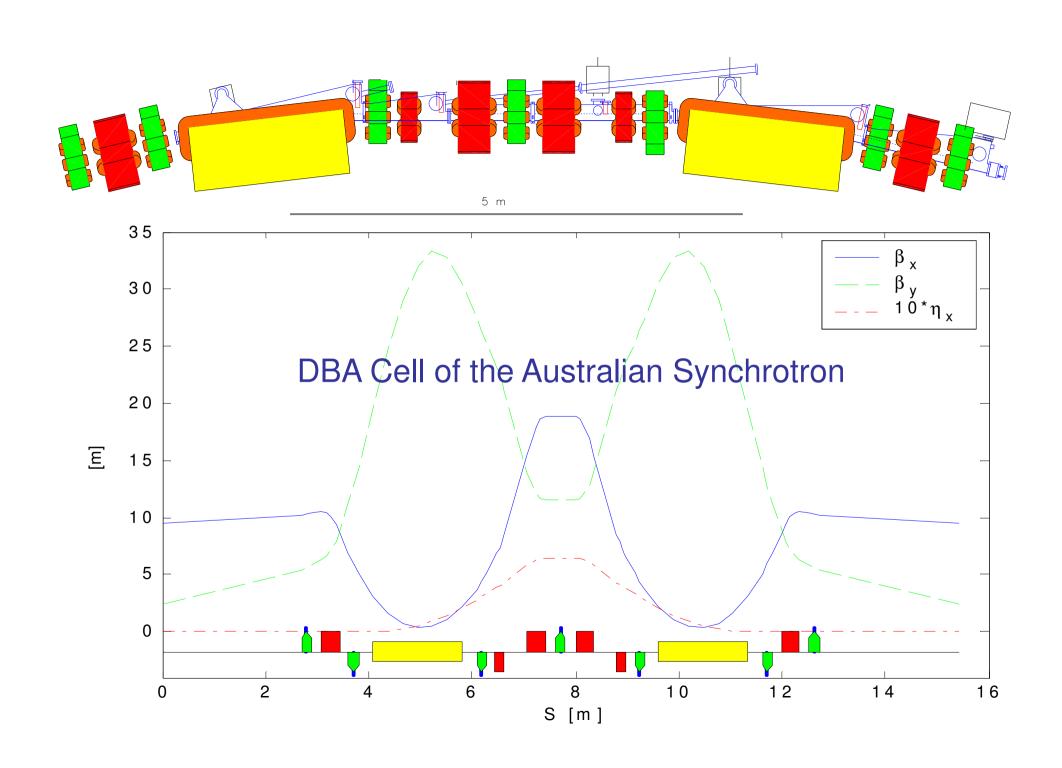
What can be done to make $\epsilon_{\bf x} \propto {f E^2} \cdot {f heta^3} \cdot {f \Gamma}$ small ?

Large ring with lots of cells, reduce electron beam energy



- → many cells → large storage rings
 - ESRF: 32; APS: 40; Spring-8: 48
- → High power "damping wigglers": much more SR produced





Large rings, many bending magnets, reduce energy

Wavelength corresponding to the diffraction limit: $\lambda_{\mathbf{DL}(\mathbf{x},\mathbf{y})} = 4\pi\epsilon_{\mathbf{x},\mathbf{y}}$

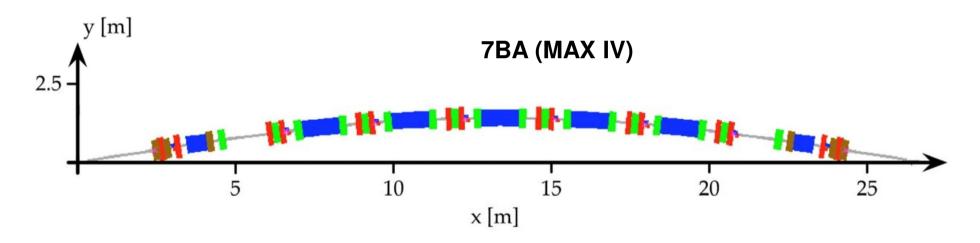
| Source | energy GeV | ϵ_x nmrad | max. B_n 10^{20} | circ. m | $\mathop{ m \AA}^{Dlx}$ | lattice |
|------------|---------------|--------------------|----------------------|------------|-------------------------|---------|
| ESRF | 6 | 4 | 3.0 | 844 | 500 | DBA |
| APS | 7 | 3.1 | 1.4 | 1104 | 390 | DBA |
| SPring-8 | 8 | 2.7 | 2.5 | 1436 | 340 | DBA |
| DIAMOND | 3 | 2.7 | 1.5 | 561 | 340 | DBA |
| PETRA III | 6 | 1 | 10 | 2408 | 126 | DBA+DW |
| NSLS II | 3 | 0.5 | 30 | 792 | 62.8 | DBA+DW |
| PETRA III* | 3 | 0.16 | ••• | 2408 | 20 | DBA+DW |



What can be done to make $~\epsilon_{f x} \propto {f E^2} \cdot heta^{f 3} \cdot \Gamma~$ small ?

Same size Ring, More Bending Magnets

"New" Idea: multi bend achromat (MBA)



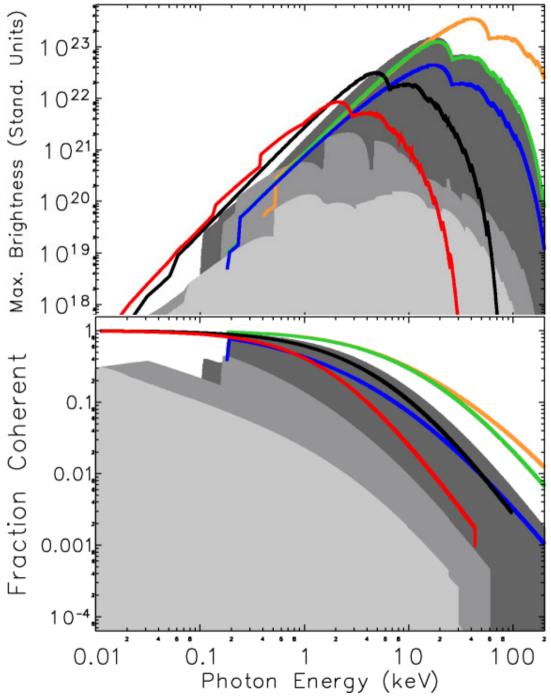
- → many small (weaker) bending magnets per cell
 - compact design
 - combined function magnets



Towards Diffraction Limited Storage Rings– Multi-Bend Achromat Lattices

Wavelength corresponding to the diffraction limit: $\lambda_{DL(\mathbf{x},\mathbf{y})} = 4\pi\epsilon_{\mathbf{x},\mathbf{y}}$

| Source | energy GeV | ϵ_x nmrad | max. B_n 10^{20} | circ. m | $ ho_{Dlx}$ Å | lattice |
|---------------|---------------|--------------------|----------------------|------------|---------------|---------|
| ESRF | 6 | 4 | 3.0 | 844 | 500 | DBA |
| APS | 7 | 3.1 | 1.4 | 1104 | 390 | DBA |
| SPring-8 | 8 | 2.7 | 2.5 | 1436 | 340 | DBA |
| DIAMOND | 3 | 2.7 | 1.5 | 561 | 340 | DBA |
| PETRA III | 6 | 1 | 10 | 2408 | 126 | DBA+DW |
| NSLS II | 3 | 0.5 | 30 | 792 | 62.8 | DBA+DW |
| PETRA III* | 3 | 0.16 | • • • | 2408 | 20 | DBA+DW |
| MAX IV | 3 | ≈ 0.25 | 40 | 528 | 31 | 7BA |
| SIRIUS | 3 | 0.28 | 20 | 518 | 35 | 5BA |
| ESRFII | 6 | 0.16 | 100 | 844 | 20 | 7BA |
| APS II | 6 | ≈ 0.07 | 200 | 1104 | 8 | (5-8)BA |
| Spring-8 II | 6 | 0.10 | 100 | 1436 | 13 | 5BA |



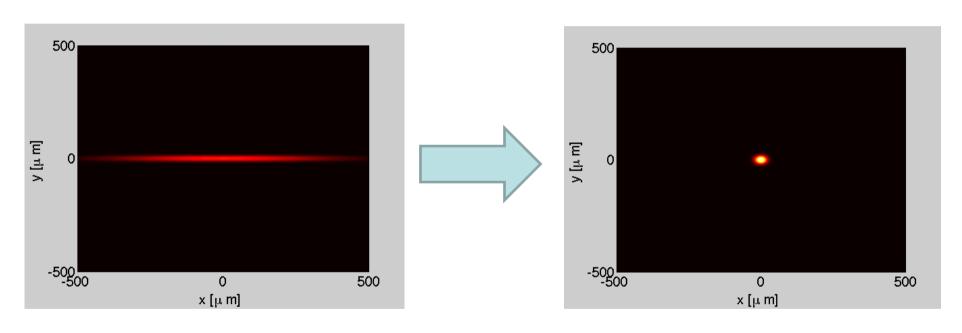
Brightness and Coherence of current & planned US SR facilities, and some Diffraction Limited Storage Ring designs.

present US facilities (light grey), US rings under construction and upgrade project (medium grey), foreign projects and plans (dark grey). diffraction-limited storage ring designs: Red: ALS-2, 0.2km/2.0GeV, 52pm, Black: NSLS-3, 0.8km/3.0GeV, 30pm, Blue:APS-2, 1.1km/6.0GeV, 80pm, Green: PEP-X, 2.2km/6.0GeV, 5pm, Orange: TAU, 6.2 km/9.0 GeV, 3 pm

Report of US BESAC Subcommittee on Future Light Sources: 2013



A "Nicer" Source Too!



Current 3rd Generation Ring

DLSR



New Storage Ring Projects

PETRA III @ DESY



Best emmittance today: $\varepsilon_h = 1 \text{ nm rad } @ 6 \text{ GeV}$

MAX IV in Lund



Under construction $\epsilon_h = 0.2\text{-}0.3 \text{ nm rad } \text{@ } 3.7 \text{ GeV}$

NSLS II @ BNL



Under construction $\epsilon_h = 0.55 \text{ nm rad } @ 3 \text{ GeV}$

Upgrades

APS@ ANL



 $\epsilon_h = 0.07 \text{ nm rad @ 6 GeV}$

ESRF in Grenoble



 $\epsilon_{h} = 0.1-0.15 \text{ nm rad @ 6 GeV}$

Spring-8 in Hyogo, Japan



 $\varepsilon_h = 0.11 \text{ nm rad } @ 6 \text{ GeV}$



Enjoy the Cheiron School!







Thanks to many people for slides Particularly Edgar Weckert, DESY

